

**Experimental Opportunities**

**~~LHC Challenges~~ for String Theory**

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**String Pheno' 14**  
**July 10, 2014**

Yale University



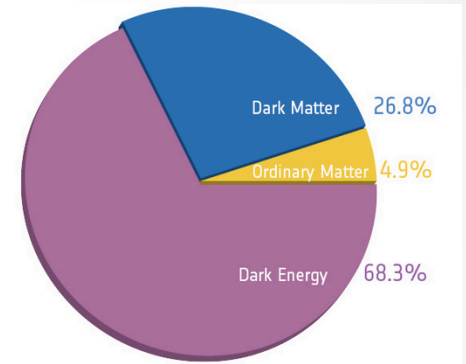
## **String Phenomenology :**

**Study of aspects of potential solutions of String Theory with features “similar” to those observed in our Universe**

# What are the “Broad Features” of Our Universe ?

- **3 + 1 “Large” dimensions ( $R_{\text{extra}} < \sim 10^{-16}$  cm)**
- **Flat Universe with Dark Energy, Dark Matter**

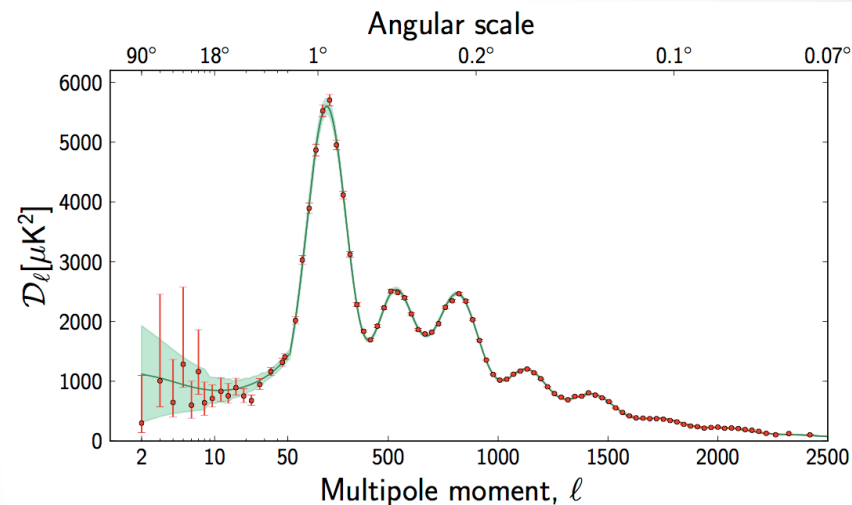
**& Ordinary \* Matter** --  $\Lambda$ CDM cosmology



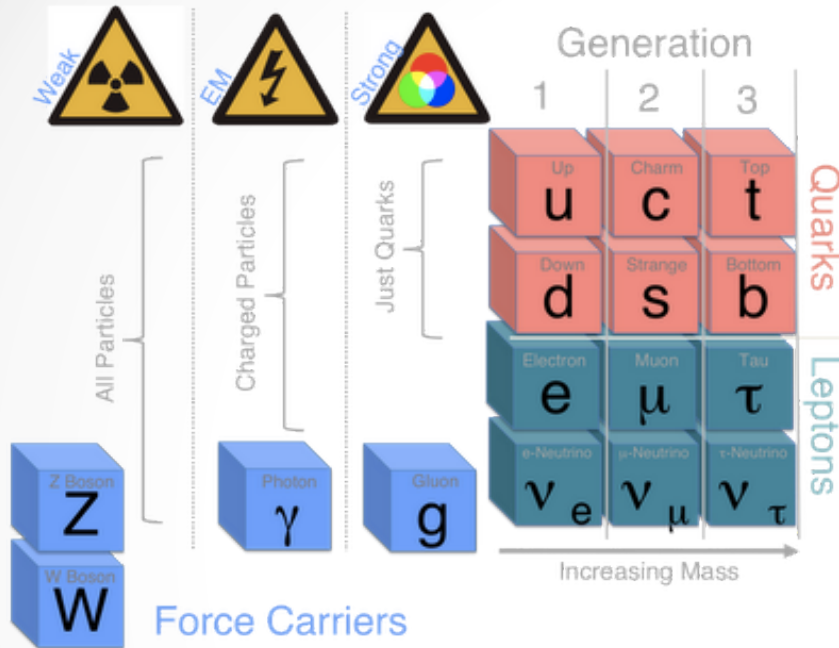
**Nearly scale invariant & adiabatic cosmological fluctuations**

**-- CMB & Large-scale structure**

**PLANCK**



# Ordinary\* Matter – Standard Model



- \* Elementary Matter (point-like)
- \* Force Carriers (gauge bosons)
- \* Scalar particle called the Higgs Boson



## Broad Features

- Non-abelian gauge theory.
- Chiral fermions
  - charged : hierarchical masses & small mixing.
  - neutral (neutrinos) : tiny, hierarchical masses & large mixing.
- Spontaneous symmetry breaking by the Higgs mechanism.

# Recent Experimental Results

## • Energy Frontier

- Positive: Discovery of Higgs @ 125 GeV.
- Null : Lack of Beyond SM physics so far

## • Intensity/Precision Frontier

- Positive : measured PMNS angle --  $\sin^2 \theta_{13}$
- Null : No deviations from SM, more stringent constraints on new physics

## • Dark Frontier

- Null : LUX (direct detection), FERMI (indirect detection),...
- Hints (?) : X-ray line, Diffuse photons from GC, ...

## • Cosmic Frontier

- Null : No sign of non-gaussianity so far
- Hint (?) : Primordial Gravitational Waves (BICEP2)

*D. Marsh, M. Rummel*

*Westphal, Burgess, Hebecker, Maharana, Takahashi, Sagnotti, Nilles, Grimm, Shiu, Kaloper, Uranga. Also many parallel talks ...*

# Plan of Talk

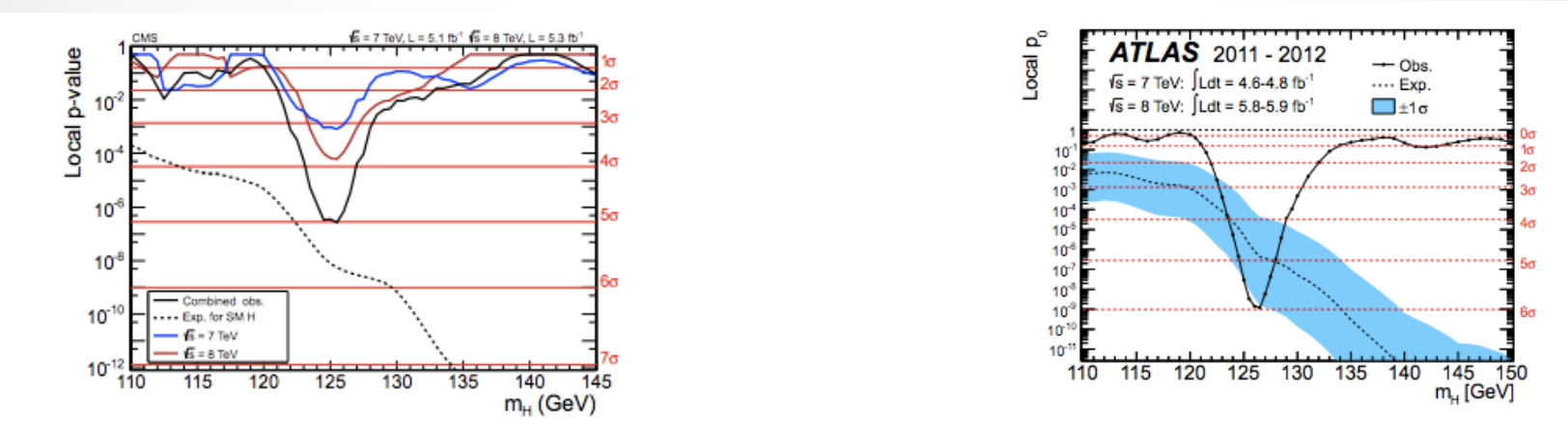
- **I) Higgs Discovery**
  - Summary of Results
  - What kinds of New Physics models favored/disfavored by data?
- **II) String motivated SUSY Models – Basic Features & Potential Signals**
  - “Electroweak-Tuned”
    - “Imperfectly” Natural
    - “Mostly” Un-natural
  - “Electroweak-Natural” – briefly discuss one possibility
  - R-parity or Not ?
- **III) Dark Matter – motivated from String Theory**
  - Status of LSP WIMP DM.
  - Dark sectors.
- **IV) Summary & concluding remarks**

# **I) Higgs Discovery**

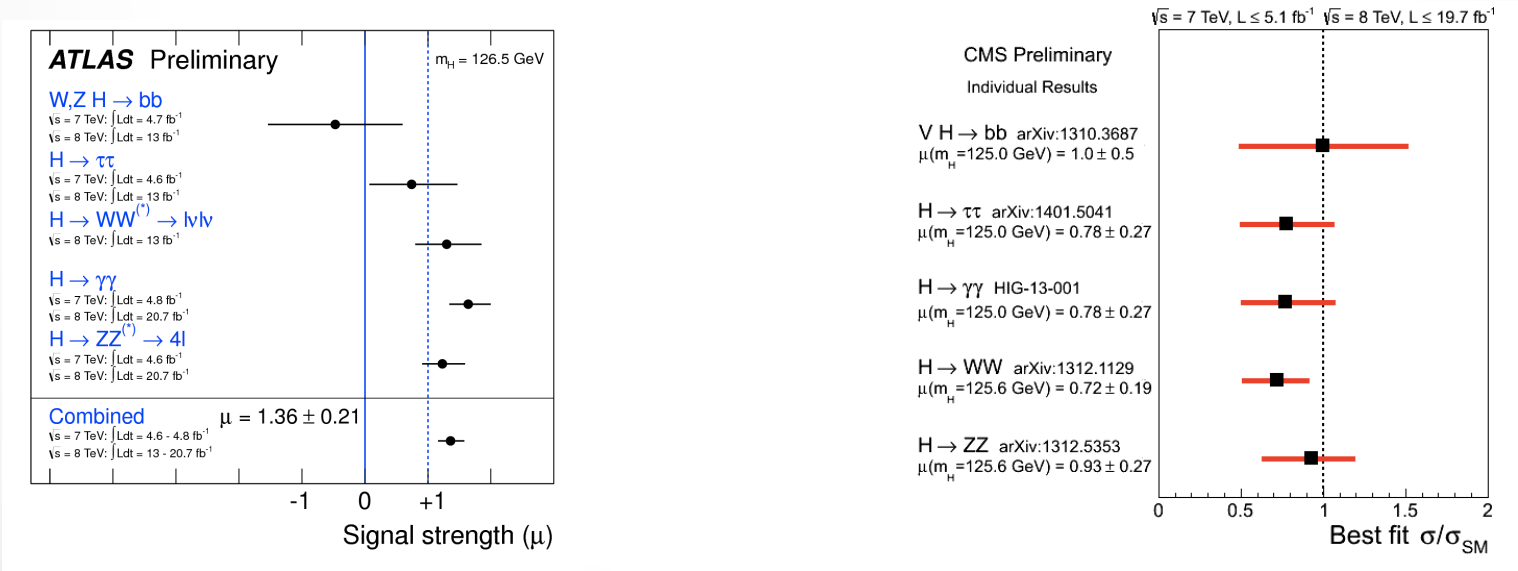
**-- Summary of Results**

**-- Effect on BSM Models**

- Discovery of a Higgs Particle @ 125 GeV



- Signal consistent with that of a SM-like Higgs



# Effect on BSM Physics

- i) Technicolor Models

- $v_{EW} = f$  (scale of New Strong Dynamics)
- All new resonances at  $M \sim 4 \pi f = 4 \pi v_{EW}$

SM-like “Higgs” resonance at  $M < \sim v_{EW}$   Strongly Disfavored

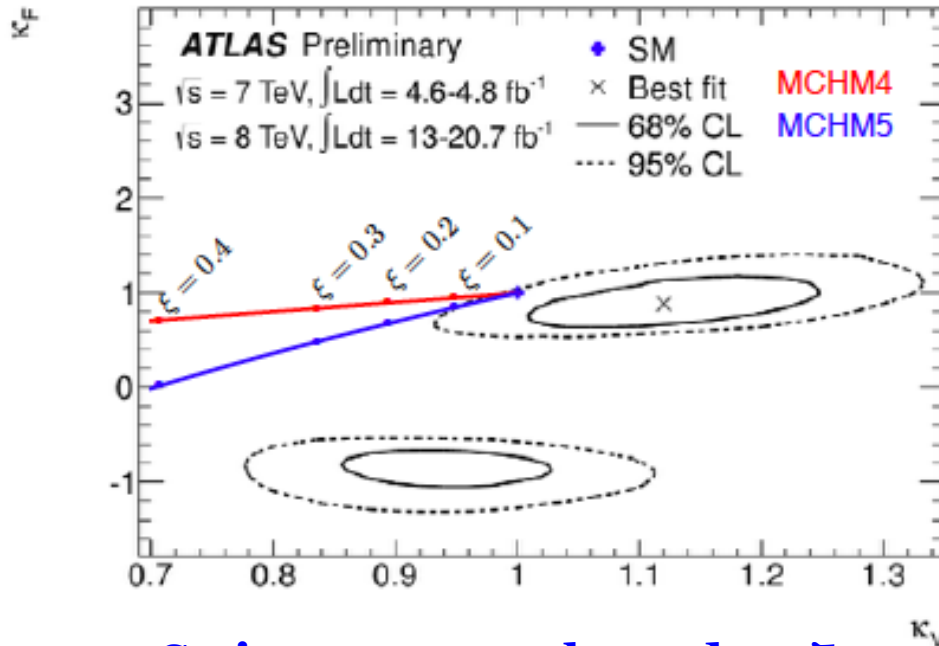
- ii) Composite Higgs Models

- Higgs pseudo-NGB of global symmetry (E.g.  $SO(5)/SO(4)$ )
- Could be light compared to  $4 \pi f$
- EW Precision observables under control, if  $\xi = v_{EW}/f < 1$ .

However, composite Higgs models in tension with data

## Composite Higgs Models (contd.)

- Higgs couplings different from that in SM at tree-level  
-- The parameter  $\xi = v_{EW}/f$  controls the deviation.
- Also expect “Top-partners” below  $\sim 1$  TeV (None observed)



$\kappa_F \sim$  Coupling to Fermions

$\kappa_V \sim$  Coupling to Gauge bosons

Same with Randall-Sundrum Models, since these are “duals” of composite Higgs

Stringent upper bound on  $\xi$

Still possible, but increasingly disfavored

### iii) TeV-scale Strings

*Antoniadis; Hashi, Wan-Zhe*

- Well-known that String scale can be made very small ( $> \sim \text{TeV}$ ) at the expense of making extra dimensions very large (relative to  $M_s$ )

$$M_P^2 = \frac{1}{g_s^2} M_s^{2+n} R_\perp^n,$$

*Arkani-Hamed, Dvali, Dimopoulos hep-ph/9803315  
Antoniadis et al hep-ph/9804398*

- **Experimental Signals**

- Kaluza-Klein Excitations

- String Resonances with Regge behavior :  $M_n^2 = nM^2, \quad j = j_0 + \alpha' M_n^2$

- Production of Black-Holes

$M_{\text{BH}} \sim M_s/g_s^2$ , **so threshold higher than that for string resonances.**

- Z' bosons with mass  $M_{Z'} \sim g M_s$  generic in isotropic compactifications.

**receive mass by Green-Schwarz mechanism : Stuckelberg U(1)**

# Constraints & Interpretation

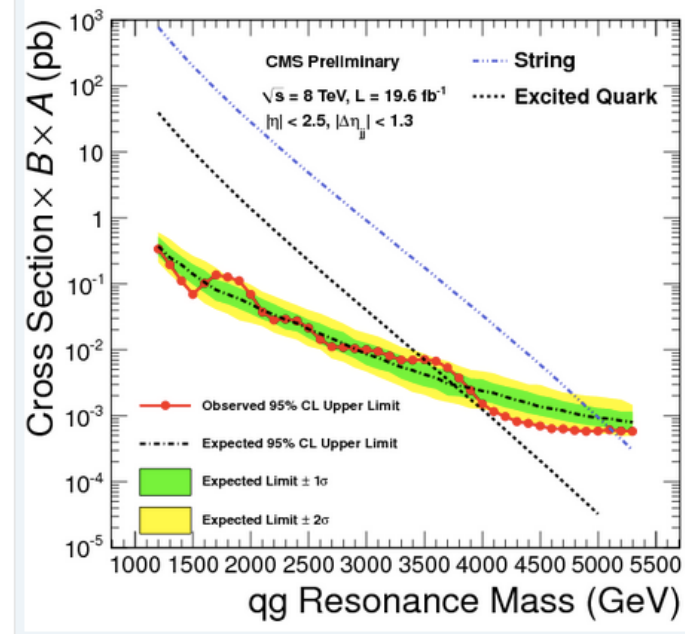
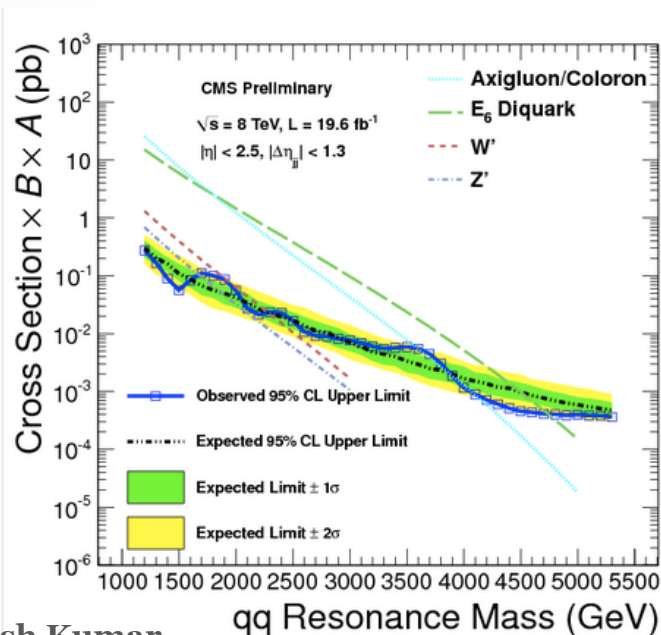
- String resonances produced in  $qg \rightarrow qg$ ,  $gg \rightarrow gg$  scattering  
-- universal amplitude:

*Anchordoqui, Goldberg, Lüst, Nawata, Steiberger, Taylor 0808.0497*

- current bound  $M_s > \sim 5$  TeV

*Also see Lüst, Taylor 1308.1619*

- FCNCs impose stronger bounds generically
- Also, bounds on  $Z'$  (for isotropic)



$M_{Z'} > \sim 2$  TeV if  $Z'$  couples the same as  $Z$

## • Effect of Higgs Discovery @ $M_H$ near 125 GeV

- EWSB can occur with  $M_H$  suppressed relative to  $M_s$  by loop factor

*Antoniadis, Benakli, Quiros NPB 583 (2000) 35*

*Antoniadis, Dimopoulos, Pomarol, Quiros NPB 544 (1999) 503*

  
>~ 5 TeV

\* So,  $M_H$  in the correct range, **however at tree-level :  $M_H = M_Z$**

\* Need large corrections to Higgs Quartic  $\lambda$  to raise  $M_H$  to 125

GeV, from KK and string modes (**not clear if fully computable**)

**Prospect :** Does not seem likely, but may still be a possibility

## II) (String-motivated) SUSY Models

(with a high  $M_s$ )

# SUSY Models


- Favored by current Experimental Data over other approaches.

## Reasons:

a) contain a Higgs-like boson with mass  $M_h < \sim v_{EW}$

b) possess a “decoupling limit” :

when  $M_{soft} \gg v_{EW}$ , Higgs SM-like & superpartners heavy

- Of course, string theory  SUSY at microscopic level
- However, nature of favored SUSY models different from naïve expectations ....

## At Crossroads.....



- **Electroweak-Tuned**  
but “minimal”

- Seems more likely
- Talk more about it

**Higgs @ ~125 GeV**

**No superpartners so far**

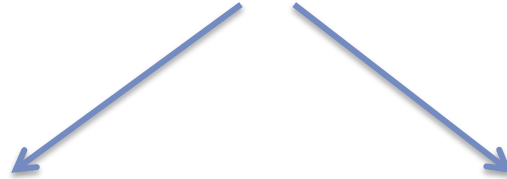
- **Electroweak-Natural**  
but “elaborate”

- Possible, but  
challenging
- Briefly mention one  
possibility

**More Data will Decide !**

# Electroweak-Tuned

- Higgs Mass @ 125 GeV points roughly to two sub-classes



## A) “Imperfectly” Natural

- \* Scalar superpartners  $M_{\text{soft}} - O(10-100)$  TeV
- \* Gauginos may be naturally suppressed by (moduli) dynamics
- \* Can explain **most** of the Hierarchy
- \* An unexplained Little Hierarchy remains

## B) “Mostly” Un-natural

- \*  $\lambda_H = 0$  at  $M_{\text{soft}} \sim O(10^{10})$  GeV
- \* Gauginos may or may not be suppressed relative to  $M_{\text{soft}}$
- \* Most of the Hierarchy is **NOT** explained – invoke fine-tuning

Talk about both :

a) Basic characteristics b) Potential Experimental Signals

## A) “Imperfectly” Natural

- **Question :** What sets the mass-scale of the scalar superpartners?

### **Elegant Solution :** Moduli Dynamics

For “Generic” Kahler potential for Moduli & Matter Fields,

$$M_{\text{modulus}} \sim M_{\text{soft}} \sim M_{3/2}$$

*Denef, Douglas hep-th/0411183 ; Gomez-Reino, Scrucra hep-th/0602246 ;  
Acharya, Kane, Kuflik 1006.3272*

- \* Can be obtained from Theory for  $O(1)$  choices of microscopic constants

*Acharya, Bobkov, Kane, **PK**, Shao PRD 76 2007, 126010*

$$M_{3/2} = O(10-100) \text{ TeV}$$

- \* Moduli heavy enough to decay before BBN.

- \* Higgs mass can be successfully computed

*Kane, **PK**, Lu, Zheng PRD 85 2012, 075026*

- **If  $H_{\text{inf}} > M_{3/2}$ , then Moduli dominated Universe before BBN.**
    - Potentially important implications for Cosmology/Astrophysics,  
E.g. growth of substructure at small scales *Erickcek, Sigurdson 1106.0536*  
*Fan, Ozsoy, Watson 1405.7373*
    - Crucial implications for Dark Matter in terms of candidates, abundance, interactions
- One example : “Non-thermal WIMP Miracle”** Talk about DM later
- Moroi, Randall hep-ph/9906527; Acharya, Kane, PK, Watson 0908.2430,*  
*Many follow-up works in the literature*

- **What about gaugino masses?**

Gaugino masses naturally suppressed relative to scalars in many string frameworks:

*Choi, Falkowski, Nilles, Olechowski, Pokorski hep-th/0411066; hep-th/0503216*  
*Conlon, Quevedo hep-th/0605141; Acharya, Bobkov, Kane, PK, Shao hep-th/0701034*

## **A i) Collider Phenomenology of Framework with “Heavy” scalars & “Light” Gauginos**

- Broad features applicable to all models in this framework.
- Precise constraints & signals depend on particular models.

# Constraints & Prospects @ LHC

- Since light(er) particles – chargino, neutralino, gluino

## -- Main Production Processes at the LHC:

$$pp \rightarrow \tilde{g}\tilde{g} \quad \chi_2^0\chi_1^\pm \quad \chi_1^\pm\chi_1^\pm$$

Strong

Electroweak

Other channels, such as  $\chi_1^0\chi_2^0$ ,  $\chi_1^0\chi_1^0$ ,  $\chi_2^0\chi_2^0$  more model dependent

## -- Decays at the LHC:

$$\text{gluino} \longrightarrow \chi_1^0 t\bar{t} \quad \chi_2^0 t\bar{t} \quad \chi_1^\pm b\bar{t}, \quad t\bar{b} \quad \chi_2^0 b\bar{b} \quad \chi_1^0 b\bar{b}$$

$$\chi_2^0 \longrightarrow \chi_1^0 Z; \quad \chi_1^0 h, \dots$$

$$\chi_1^\pm \longrightarrow \chi_1^0 W^\pm; \dots$$

Again, precise BR's model-dependent

**Typical Final State: High  $p_T$  multi-jets,  $\geq 3$  b jets + 0 or 1 lepton +  $E_T$**

# Constraints and Future Prospects

- **Analysis of LHC data presented in terms of “simplified models”**
  - assume 100% BR of  $\tilde{g}$  to one channel, for e.g.  $\tilde{g} \rightarrow \chi_1^0 t \bar{t}$
  - Real Models  $\longrightarrow$  BR's to many channels can be non-trivial.
  - Should compute bound on masses for each model.

## E.g. 1 : **String-Motivated SO(10) Yukawa Unification Models**

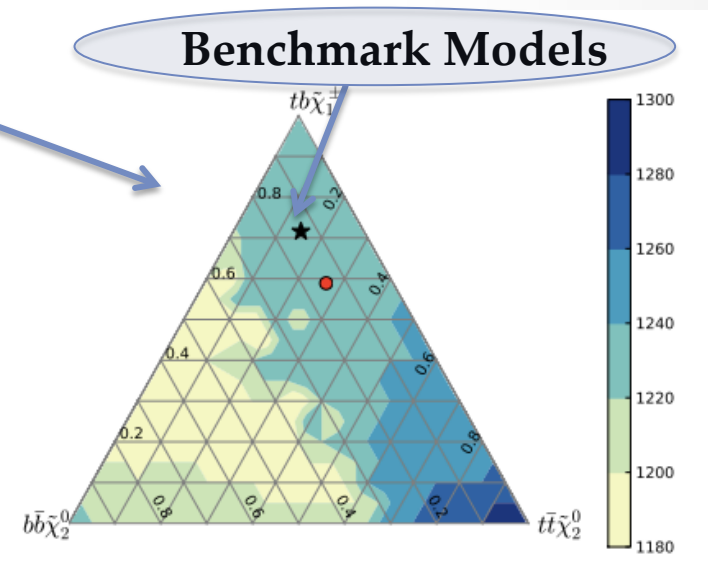
**Bounds on Gluino Masses :** *Raby 1309.3247*

*Anandakrishnan, Bryant, Raby 1404.5628, 1303.5125*

## E.g. 2 : **M-theory motivated G<sub>2</sub>-MSSM**

*S. Ellis, Kane : To be published*

**Similar technique :**  $M_{\text{gluino}} > 0.9\text{-}1 \text{ TeV}$



**Implicit Assumption: Gluino reasonably heavier than the LSP**



**jets, leptons “hard”, i.e. have large  $p_T$ .**

- **However, if Gluino and LSP close in mass, then**
  - spectra “compressed”
  - jets & leptons “soft”, many do not pass cuts.

**Eg : “Mirage Mediation Models” with Precision Gauge Unification**

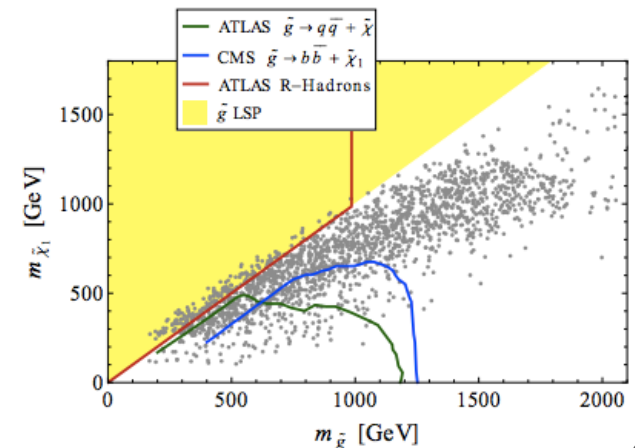
*Choi, Jeong, Kobayashi, Okumura hep-ph/0508029*

*Krippendorff, Nilles, Ratz, Winkler 1306.0574*

Also see Pheno. Papers on ‘compressed SUSY’



- Gluino bounds considerably relaxed.
- Gluino may be long-lived ( $10\mu\text{m} - 1\text{ mm}$ )
- Co-annihilation effects important for LSP annihilation in early Universe.



# Prospects @ ~100 TeV Collider

*I. Antoniadis' Talk*

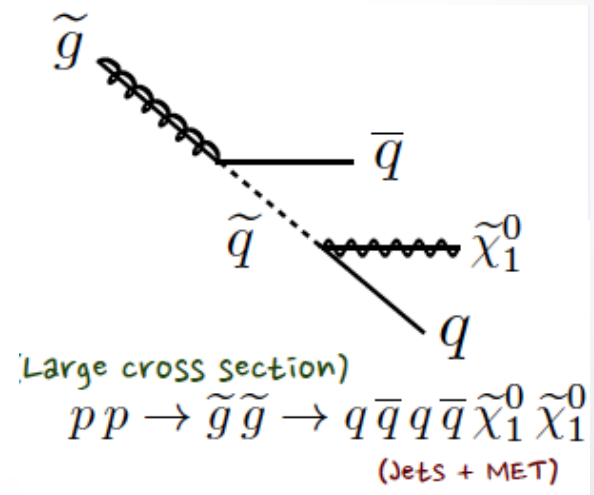
- LHC sensitive to gluino masses  $< \sim 2\text{-}2.5$  TeV.
- A larger CM-Energy Collider will increase the reach.
- Studies quite preliminary. Lot to do .
- Some work done for “simplified models”

*Cohen et al 1311.6480*

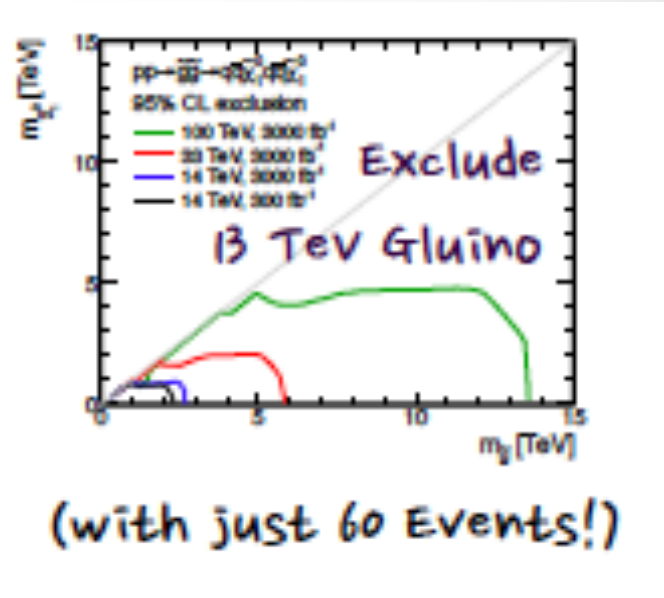
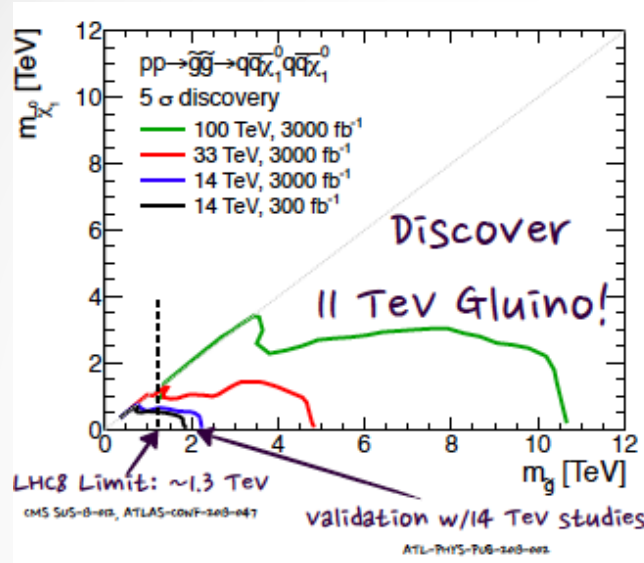
**E.g. Gluino-Neutralino Simplified Model**

**Final State : Multi-jets + No Leptons +  $E_T$**

**Dominant Background :  $t\bar{t}$  + jets**  
(in contrast to  $W/Z$ +jets @ LHC)



- **Significant improvement compared to the LHC** *Cohen et al 1311.6480*



- **What about stops & heavier electroweak-inos?**

- No detailed studies yet. *Acharya, Bozek, Pongkitivanichkul, Sakurai: To appear*
- May be possible to detect these for  $M_{3/2} < \sim 30$  TeV

\* LHC has a good chance of discovering “Imperfect-Naturalness”

\* 100 TeV collider would be a wonderful development

- would greatly help in confirming/ruling out the above and also other ideas

## B) “Mostly” Un-natural

*L. Ibanez's Talk*

### \* Basic Motivation:

SM -- well-known that certain values of Higgs Mass can be tied to vanishing of Higgs quartic  $\lambda$  at some High scale.

**Higgs near 125 GeV –  $\lambda$  vanishes at  $M \sim 10^{10}$  GeV**

(uncertain due to  $M_{\text{top}}$  uncertainty)

*Elias-Miro et al 1112.3022 ; Holthausen et al 1112.2415; Wetterich 1112.2910,...*

Most of the Hierarchy NOT explained, just fine-tuned

### • Proposal:

**SUSY @ High scale  $M_{\text{soft}}$  such that  $\lambda(M_{\text{soft}}) \rightarrow 0$  at  $M_{\text{soft}}$**

**But** 
$$\lambda(m_S) = \frac{g^2(m_S) + g'^2(m_S)}{8} \cos^2 2\beta \longrightarrow \tan \beta = 1 @ M_{\text{soft}}$$

Can be motivated from theoretical approaches:

*Hebecker, Knochel, Weigand 1204.2551; 1304.2567*

*Ibanez, Marchesano, Regalado, Valenzuela 1206.2655; Ibanez, Valenzuela 1301.5167*

- SUSY at  $M \sim 10^{10}$  GeV can be combined with Gauge Unification in F-Theory : Both scales can be related**

$$M_{SS} = ((2g_s)^{1/2} / \alpha_G^{1/2}) \frac{M_c^2}{M_p}$$

**Unification @  $M_c \sim 10^{14}$  GeV with threshold corrections.**

*Ibanez, Marchesano, Regalado, Valenzuela 1206.2655*

*Camara, Ibanez, Valenzuela 1404.0817*

- Can give rise to QCD Axion with decay constant  $F_a \sim 10^{12}$  GeV**

- However, proton decay with a low Unification scale a challenge**

*Hebecker, Unwin 1405.2930*

**“Fake Split-SUSY”** – *Goodsell’s talk.*

*Hebecker’s talk*

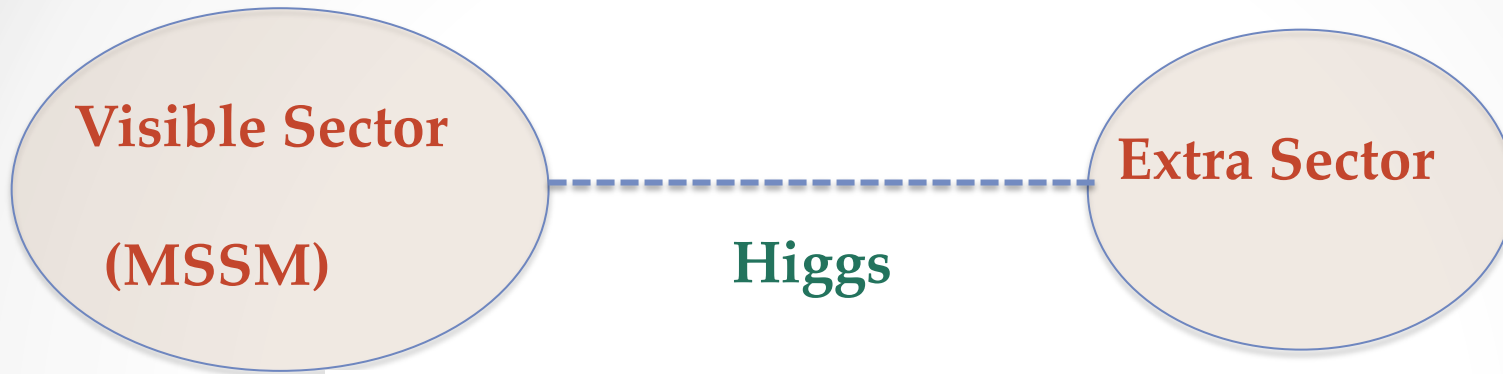
**Experimental Probes (if no light fermions)**

- Precise Measurements of Higgs & Top Mass & couplings.**
- Possible discovery of QCD Axion DM in ADMX with  $F_a \sim 10^{12}$  GeV.**
- Observation of Proton Decay.**

# **“Electroweak-Natural”**

- **As mentioned earlier, current data makes it challenging to realize this possibility.**
- **However, Nature may still work this way. Within SUSY, have to go beyond the MSSM :**
  - **Additional contributions to Physical Higgs Mass**
  - **New contributions to the Higgs potential**
    - \* **may improve naturalness of EWSB**
  - **No Beyond-the-SM physics so far → models more “elaborate”**
- **Fully explicit and viable models hard to construct. Nevertheless many attempts in literature.**
- **Talk about one possibility.**

# “Holomorphic” Higgs Portal



$$W = \lambda_u \mathcal{H}_u \mathcal{O}_u + \lambda_d \mathcal{H}_d \mathcal{O}_d.$$

Higgs couples to operators in the Extra sector (in the superpotential).

$\mathcal{O}_u, \mathcal{O}_d$  -- part of SUSY breaking sector or part of messenger sector which couples to another SUSY sector -- **extension of Gauge Mediation to Higgs sector.**

**Such terms considered in various field-theoretic contexts ...**

*Azatov et al 1106.3646, 1106.4815; Kitano et al 1206.4053; Stancato et al 0807.3961  
Gherghetta, Pomarol 1107.4697; Komargodski, Seiberg 0812.3900; Craig et al 1302.2642,  
Knapen et al 1311.7107, Schafer-Nameki et al 1005.0841, ....*

- Setup could arise naturally in a class of string frameworks**

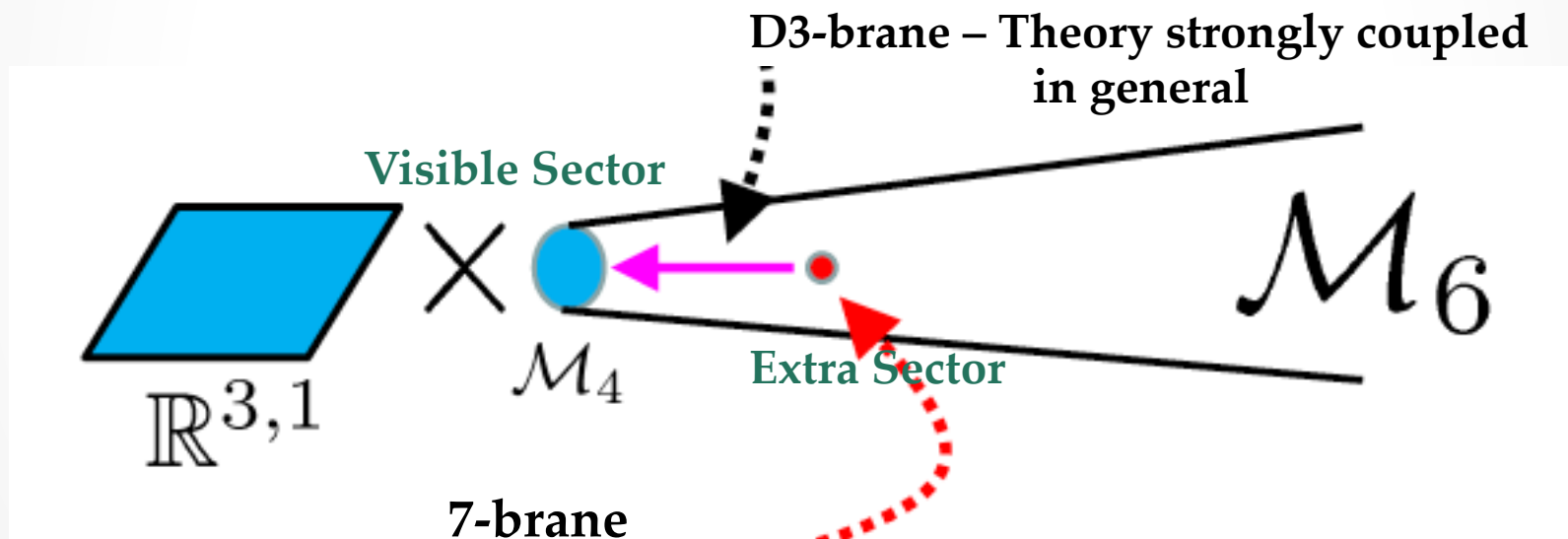
**“F-Theory” --- 7-brane probed by a D3-brane**

*Heckman, Vafa 1006.5459*

*Follow-ups with collaborators*

Visible Sector

Extra Sector



**Local Model -- study region in which D3-brane is close to the 7-brane**

# Phenomenologically interesting Features

-- Higgs Potential could change relative to the MSSM

*Heckman, PK, Vafa, Wecht JHEP 1201 (2012)*

-- Consistent with gauge coupling unification in the MSSM.

*Heckman, Vafa, Wecht 1103.3287*

-- Possible to compute Higgs couplings : (Using SUSY, Holomorphy & Gauge invariance)

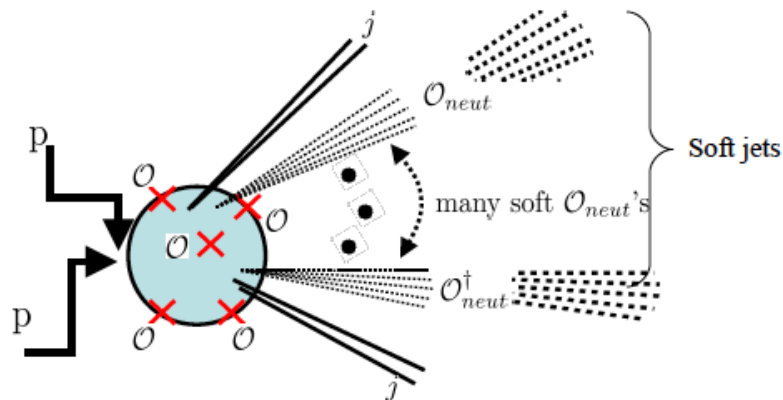
*Heckman, PK, Wecht 1204.3640; Heckman, PK, Wecht 1212.2979*

**MSSM coupled to sector which is superconformal in the UV :**

-- Imagine conformal symmetry broken with a “mass-gap”  $M$  and SUSY at scale  $(F)^{1/2}$ .

-- very interesting to understand this dynamically.

**Possible Collider  
Signal :**



**interesting  
to explore ...**

**RPV**

**RPV**

**III) To ~~be~~, or not to ~~be~~, that is the question...**

- **Until now, implicitly assumed R-parity conservation.**
- **However, possibility of R-parity violation quite interesting:**
  - **LSP no-longer stable.**
  - **Significant reduction in missing  $E_T$  @ LHC**
    - ↳ **constraints on superpartners weakened.**
  - **Viable RPV models can be constructed phenomenologically**

- **What about R-parity violation from top-down point of view?**

**Talk about : i)  $SU(5)$  GUT models, ii)  $SO(10)$  GUT models.**

Any RPV disfavored

Spontaneous RPV a possibility

*B. Ovrut's Talk*

- SU(5) GUTs:** appealing due to simplicity

- GUT breaking to  $G_{SM}$  and doublet-triplet splitting.

- employ some global symmetry  $H'$  arising in string theory

- To solve  $\mu/B\mu$  problem, either by KN/CM or GM mechanism.

*Kim, Nilles PLB138 (1984) 158; Casas, Munoz hep-ph/9302227; Giudice, Masiero PLB206 (1988) 480*

$H'$  forbids  $\mu$  parameter at High scale, but  $H'$  must be broken to  $H \subset H'$

**True in both Heterotic orbifolds & M-theory constructions**

*Kappl, Nilles, Ramos-Sanchez, Ratz, Schmidt-Hoberg 0812.2120;*

*Lee, Raby, Ratz, Ross, Schieren 1009.0905; 1102.3595;*

*Chen, Ratz, Staudt, Vaudrevange 1206.5375; Witten hep-ph/0201018;*

*Acharya, Kane, Kuflik, Lu 1102.0556*

*M. Ratz's Talk*

**Then, can show that bilinear RPV coefficient  $\kappa$  in  $\int d^2\theta \kappa L H_u$  is such that**

either a)  $\kappa/\mu = O(1)$  ( $H$  is trivial), or b)  $\kappa/\mu = 0$  ( $H$  equivalent to R-parity)

*Acharya, Kane, PK, Lu, Zheng 1403.4948*

**But stringent constraints on bilinear RPV from neutrino masses :  $\kappa/\mu < \sim 10^{-3}$**

**R-parity violation disfavored**

**Any observation of R-parity violation → disfavor above class of Models**

- **SO(10) GUTs :**

-- appealing, since 16 of SO(10) contains all SM particles + RH neutrino.

Eg :  $E_8 \rightarrow SO(10) \rightarrow G_{SM} * U(1)_{B-L}$

**Heterotic M-theory : “Exact” MSSM spectrum -- Minimal**

*Braun, He, Ovrut, Pantev hep-th/0501070; hep-th/0512177; hep-th/0602173*

- $U(1)_{B-L}$  must be broken to make  $Z_{B-L}$  sufficiently massive.

- **Since only candidate  $\langle \tilde{\nu}^c \rangle$  has odd B-L,**

→ **R-parity, a  $Z_2$ - even subgroup of B-L, is spontaneously broken**

For  $\langle \tilde{\nu}^c \rangle$  to obtain a pheno. viable vev, need :

-- **large flavor-dependent non-universality in the sneutrino soft masses relative to that for sleptons & selectrons**

*Amboroso, Ovrut, 0910.1129; Acharya, Kane, PK, Lu, Zheng 1403.4948*

Option: Have extra 10's, 16's of SO(10)

*Acharya et al :  
To appear*

# Broad Experimental Signals

*B. Ovrut's talk*

- $Z_{B-L}$  gauge boson with mass  $> \sim$  few TeV
- Existence of two light RH neutrinos.
- Leptonic RPV through the  $L H_u$  operator



**“LSP” can decay. Also, “LSP” can be charged or colored.**

- **Neutrino-Neutralino Mixing**  
-- generate majorana neutrino masses at tree level.
- **Can also have correlation between LSP decays & Neutrino Hierarchy !**

*Marshall, Ovrut, Purves, Spinner 1401.7989, 1402.5434*

**More details/signals should be explored....**

# III) Dark Matter

(motivated from String Theory)

*\* B. Dutta's Talk*

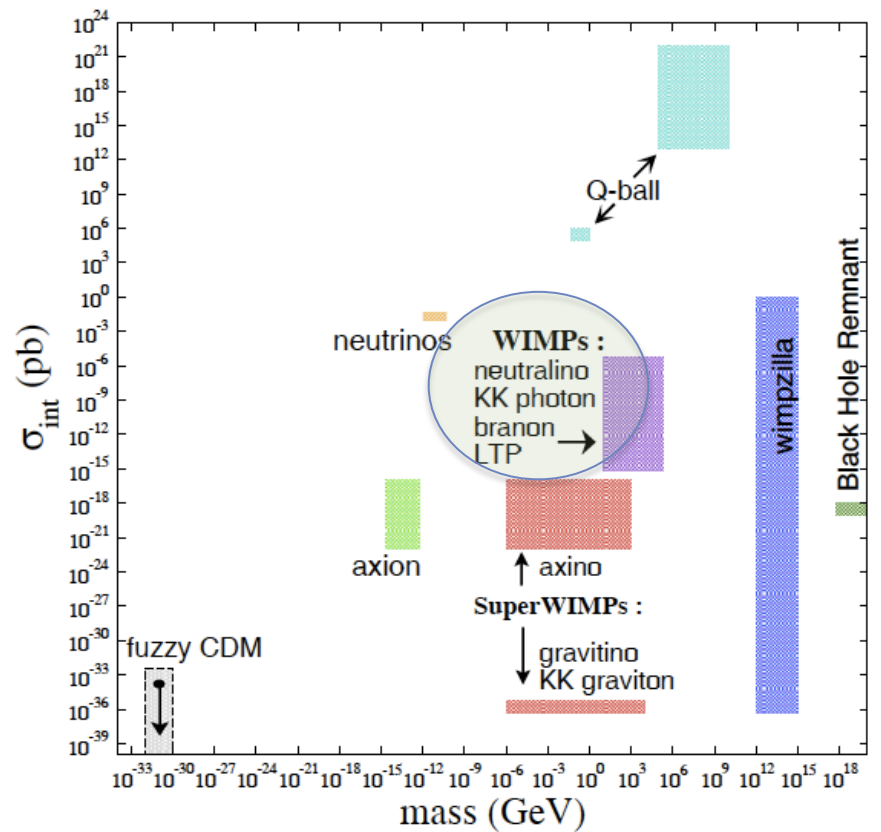
# The Dark Matter Zoo

Just an illustration – many more candidates possible ...

**Most popular candidates –**

**WIMPs & Axions**

- **String Axions – many talks**
  - could be important during inflation.
  - could also naturally comprise Dark Matter.



*Gardner et al 1303.4758*

**Proposal to detect QCD axion with GUT scale  $F_a$**   
*Graham, Rajendran 1306.6088*

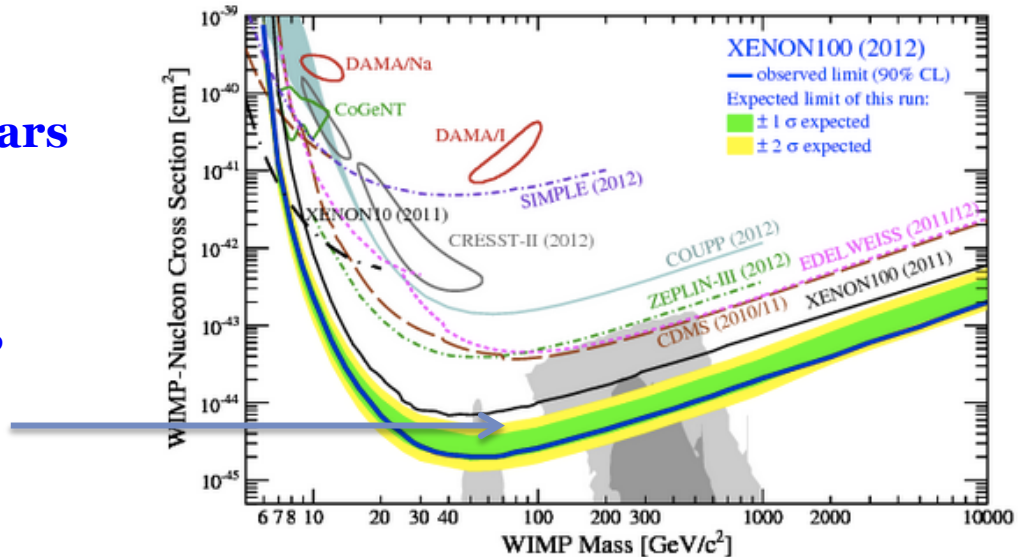
**Finally, Axions can be Dark Radiation**  
*Talks by D. Marsh, Angus, Pongkitvanichkul*

*Arvanitaki et al 0905.4720*  
*Acharya, Bobkov, PK 1004.5138*  
*Cicoli et al 1206.0819; Arias et al 1201.5902*  
*Allahverdi et al 1401.4364; Honecker et al 1312.4517, Sik Jeong et al 1310.1774;*  
*Many others..*

## (SUSY) WIMPs – minimal, since part of BSM Model.

### Direct Detection

- Many hints in the past few years
- All of them killed by LUX
- A large chunk of SUSY WIMP parameter space ruled out, and large chunks still left..



Example of SUSY WIMP not ruled out by direct detection – **Wino LSP**,

- Winos do not interact via Z-exchange or Higgs-exchange at tree level.
- Winos can also give rise to the correct abundance via the “non-thermal WIMP miracle”

**However, ...**

## Indirect detection

*Fan et al 1307.4400*

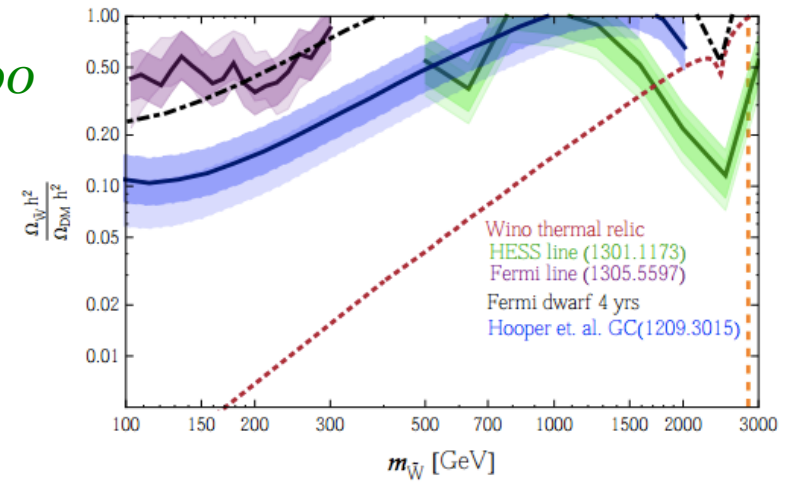
**Latest bounds disfavor Wino DM.**

**E.g. FERMI diffuse  $\gamma$  from Galactic Center**

**Also, recent hints for WIMP  
indirect detection less convincing now**

\* 130 GeV “ $\gamma$ -line” from GC

\* PAMELA Antipositron fraction from nearby region of Milky way.



**Although LSP-WIMPs still viable,**

- Constraints more & more stringent.
- In some sense, “Lamp-post” Physics.

**Worth considering other approaches**




# Dark Sectors

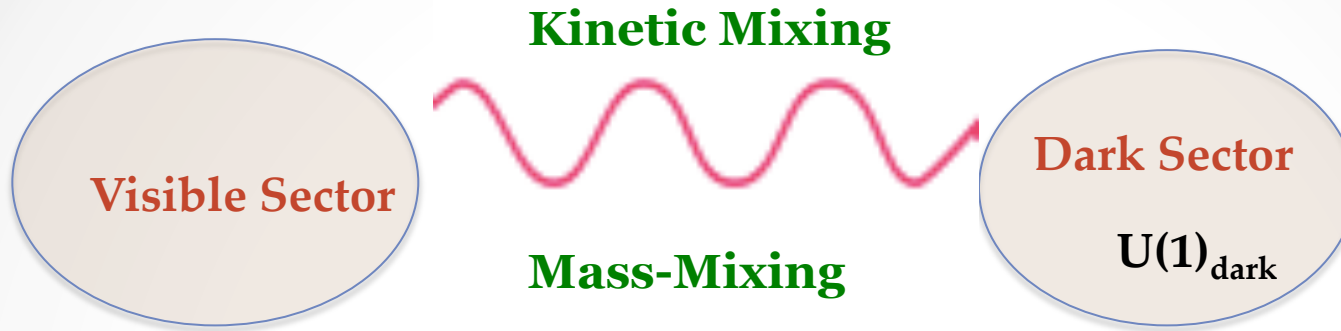
## Motivation :

- **Additional sectors in String Theory very common/natural.**
  - UV completion of SM has additional gauge/matter spectra in most cases.
  - String-consistency conditions “demand” it.
    - E.g. Hidden  $E_8$  in Heterotic, RR-Tadpole cancellation in Type II.

Dark Matter could naturally be part of these additional sectors.

- **Some “common” observations :**
  - Many talks on massive  $U(1)$ s in string theory*  

  - Extra  $U(1)$  gauge bosons --  $Z'$  (massive) ,  $\gamma'$  (massless)
    - a)  $Z'$  – Stuckelberg; b)  $Z'$  –Higgs; c)  $\gamma'$  - massless
  - Hidden sector DM or “Light” Messenger DM
    - Cvetic, Halverson, Piragua 1210.5245; Feng, Shiu, Soler, Ye 1401.5880, 1401.5890; Halverson, Orlofsky, Pierce 1403.1592; Many others....*

# Portals



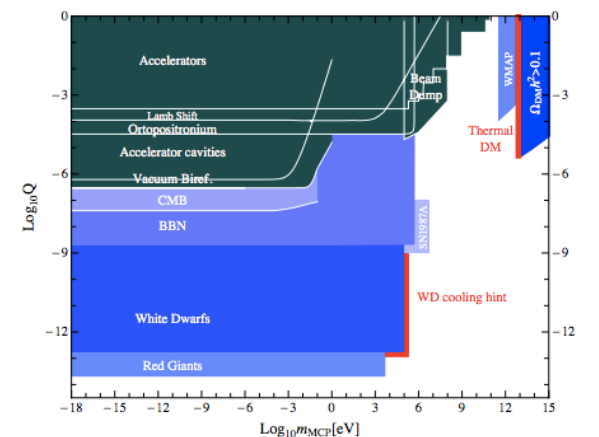
*Talks by Marchesano, Ramos-Sanchez, Mehta*

**Kinetic Mixing :**  $L \supset \int d^2\theta \epsilon W_Y W_X + h.c.$

- **Marginal coupling**  $\longrightarrow$  If generated, will persist to low energies
- **Phenomenology depends on  $\{M_{A'}, \epsilon\}$  &  $\{M_X\}$**

**a)  $M_{A'} = 0$ , Hidden sector fields acquire milli-charge  $\sim \epsilon$ .**

*Holdom PLB 166 (1986); Banks, Seiberg 1011.5120;  
Abel, Schofeld hep-th/0311051, Marchesano et al 1406.27;*



**b)  $M_{A'} \neq 0$ . Variety of  $\{M_{A'}, \epsilon\}$  can be generated.**

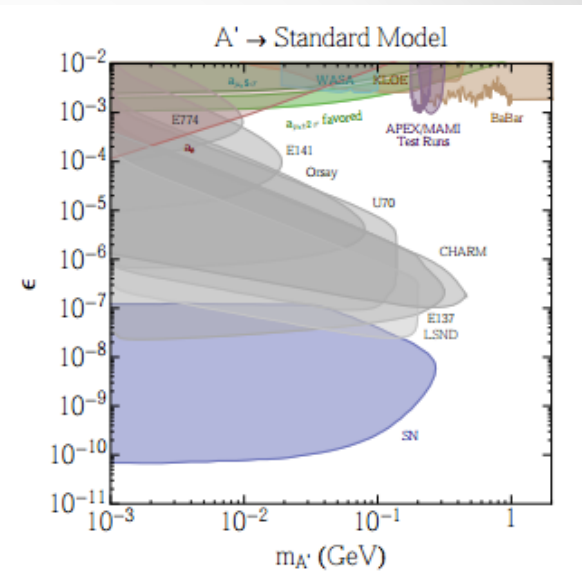
*Abel et al hep-ph/0608248, 0803.1449;  
Goodsell et al 0909.0515; 11110.6901; Cicoli et al 1103.3704*

**Dark Gauge boson will have small  
coupling  $\sim \epsilon$  to visible sector & vice-versa**

**Interesting Consequence:**

**“LSP” will decay to Dark Sector before BBN  
even with R-parity conservation.**

**E.g. Bounds on Winos can be evaded**



**Mass Mixing :**

-- **Physical  $Z'$  eigenstates: generically couple with  $O(1)$  strength to SM fermions**

*Feng, Shiu, Soler, Ye 1401.5880, 1401.5890*

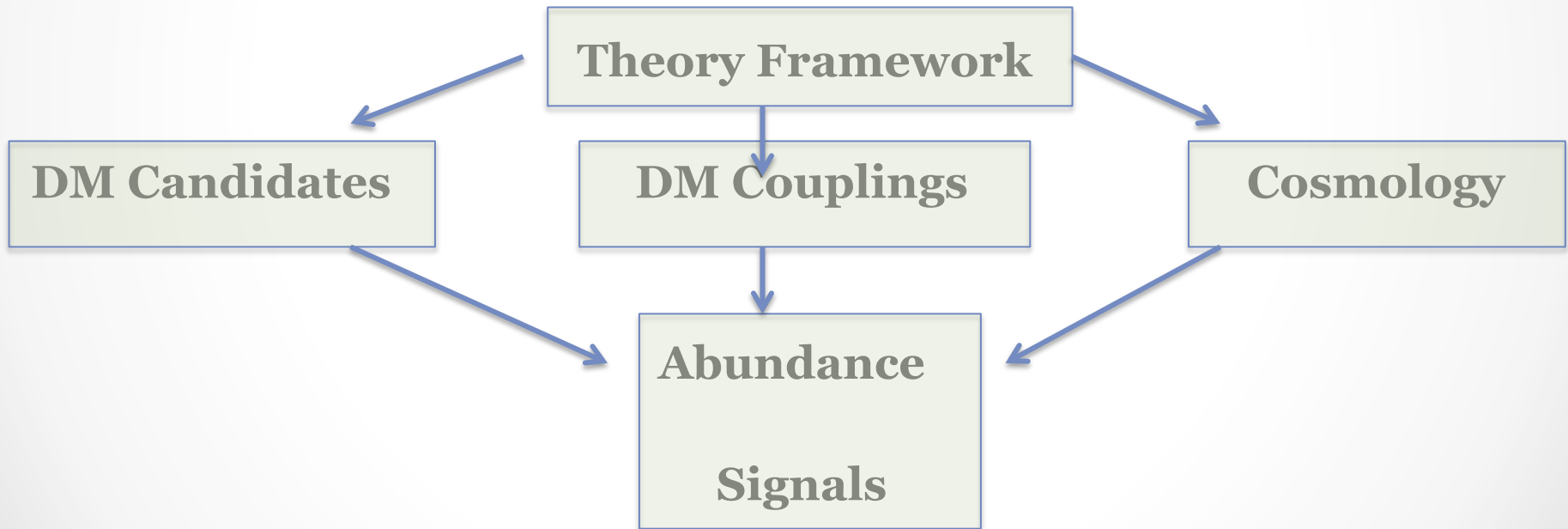
--  $M_{Z'}$  can only be suppressed by a few orders of magnitude relative to  $M_{\text{string}}$ .

-- Phenomenologically relevant only for low string scale.

-- Bound on  $M_{Z'} > \sim \text{few TeV}$

## Only the tip of the DM Iceberg...

- Until now, only talked about  $U(1)_{\text{dark}}$
- Many other possibilities :  $G_{\text{dark}}$  ;  $G_{\text{dark}} * U(1)_{\text{dark}}$  ;  $G_{\text{dark}} * G_{\text{flavor}}$  ; ....
- Important and useful to have well-motivated theoretical guide



Tons to explore ...

# Summary & Concluding Remarks

- **We are living in a data-rich era.**
  - Data, even if “Null”, can provide important insights.
- **Talked about some aspects of recent data in High-energy physics, and the insights it provides for string-motivated frameworks vis-à-vis :**
  - **Higgs and Beyond-SM physics.**
  - **Dark Matter Physics.**
- **SUSY still the most probable framework for Beyond-SM physics.**
  - **However, SUSY models different from what naively expected.**
  - **Most “simple” models appear to be “electroweak-tuned”**



-- Studied potential signals of each

- **R-parity violation** – interesting implications for string-GUT models.
  - \* **Observation of RPV will disfavor SU(5) GUT models.**  
(with mild assumptions)
  - \* **SO(10) models compatible with *spontaneous* RPV in principle**  
-- can give rise to interesting signals.
- **Dark Matter** -- Variety of possibilities
  - \* **Status of LSP WIMP DM** -- still viable but under increasing strain.
  - \* **Worth looking at other frameworks:**  
**E.g. Dark Sectors very well motivated**
    - incredible array of possibilities, just scratched the tip ...
    - very important to have an underlying theoretical framework for understanding different aspects in a coherent manner.

- **“Electroweak-Natural” Models seem rather challenging.**
  - Should not give up hope, however. May still be possible...

**Think outside the Box !**

- **Hope that Nature is kind to us and provides us with opportunities to make String Theory an experimental science.**